CO2 AS A VOLATILE COMPONENT OF THE MANTLE

David H. Eggler, Geophysical Laboratory, 2801 Upton St., N.W., Washington, D.C. 20008

There appears to be wide agreement among petrologists that the volatiles CO2 and H2O are important components of kimberlite. There is less agreement on the relative proportions of H2O and CO2 and on the role of the volatiles -- whether they occurred primarily in a low-density vapor phase, or in immiscible carbonatite melt (Koster van Groos and Wyllie, 1973), or dissolved in silicate melt, either a melt of kimberlite affinity (Dawson, 1971) or a melt encountered by the kimberlite on its eruptive path (Boyd and Nixon, 1973). This study is a determination of phase equilibria in the system Mg_SiO4-SiO2-H20-CO2 at 20 kbar. The system is considered to be a simple model of a peridotitic mantle, as it contains two of the four principal peridotite phases, forsterite (Fo, Mg₂SiO₄) and enstatite (En, MgSiO₃), plus two volatile components. The study provides information on three questions, the solubility of CO2 in silicate melt, the role of CO2 in solution of silicates in vapor, and the role of CO2 in determining the composition of magmas produced by melting of peridotite.

<u>Solubility of CO₂ in silicate melt</u>. Experiments have been previously conducted on the solubility of CO₂ in silicate melts, at 20 kbar, with the compositions CaMgSi₂O₆, MgSiO₃, and NaAlSi₃O₈ (Eggler, 1973). Solubilities ranged from 9.5 wt % (34 mole %) in relatively basic CaMgSi₂O₆ melt to only 0.9 wt % (5 mole %) in relatively acidic NaAlSi₃O₈ melt. Melts containing H₂O in addition to CO₂ dissolve slightly more CO₂. At pressures at least to 30 kbar, a large miscibility gap exists between silicate melts and H₂O-CO₂ vapor (see Fig. 1). For all compositions, CO₂ is much less soluble than H₂O; the vapor-absent region is much smaller for compositions containing more CO₂ (see, for example, relations involving the four-phase tetrahedron Fo-En-L-V at 1450° C (Fig. 1)). Melt of MgSiO₃ composition dissolves a maximum of about 3.7 wt % CO₂ at 20 kbar. In Fig. 1 the melt region is outlined by a "caterpillar" shape and is drawn on the assumption of negligible CO₂ solution in SiO₂ melt.

<u>Vapor composition</u>. The kimberlite vapor phase is often assumed to dissolve excess alkalies (Boyd and Nixon, 1973; Koster van Groos and Wyllie, 1973). Experiments on this effect are uncompleted; however, the considerable effect of CO_2 in reducing solution of silica is shown schematically in Fig. 1 and in more detail in the subsolidus region of the system En-H₂O-CO₂ in Fig. 2. When subsolidus vapor contains more H₂O than 0.65 mole %, vapor dissolves sufficient silica in excess of stoichiometric MgSiO₃ that the assemblage is Fo+En+V. However, in more CO_2 -rich vapor, very little excess silica dissolves, producing En+V.

<u>Melting relations</u>. In the system Mg_2SiO_4 -SiO_2-H₂O, compositions of liquids in equilibrium with Fo and En can be either silica-oversaturated or silica-undersaturated, depending on whether or not enstatite melts to a liquid in equilibrium with forsterite. Kushiro, Yoder, and Nishikawa (1968) have demonstrated these relations at a pressure of 10 kbar. Eggler (1973) confirmed this behavior at 20 kbar. At 20 kbar the assemblage Fo+En+V melts to H₂O-saturated liquid at 1305^o C. At temperatures below 1400^o C, liquids in equilibrium with Fo+En are silica-oversaturated while at temperatures above 1400^o C, they are silica-undersaturated. The effect of CO₂ on this relation was studied in the join 5MgSiO₃: $4CO_2-H_2O$ (Fig. 2). Two points are of interest: first, that the solidus temperature is higher in the presence of CO_2 , and second, that the temperature at which Fo+En are no longer in equilibrium with silica-oversaturated liquid (that is, where Fo+En+L+V is no longer a stable assemblage on the join En-H_2O-CO_2) is still 1400° C. Phase relations are clarified in the semi-schematic Figure 1. In

Phase relations are clarified in the semi-schematic Figure 1. In the system Mg_2SiO_4 -SiO₂-H₂O the assemblage Fo+En is in equilibrium with



Figure 1. Phase relations (semi-schematic) in the system Fo-En-H₂O-CO₂ at 20 kbar pressure (wt %). Stippled areas outline the vapor field and the four-phase tetrahedron Fo-En-L-V.

silica-oversaturated, H₂O-rich melt at temperatures below 1400^o C, whereas above 1400^o C those phases are in equilibrium with silicaundersaturated, less H₂O-rich melt. Figure 1 shows that in the presence of CO₂ the same situation holds (the difference is that, in the presence of CO₂, the solidus of a peridotite is considerably higher in temperature than 1305^o C). At 1375^o C, the stippled tetrahedron is the four-phase volume Fo+En+L+V; the vapor contains about 45 wt % CO₂ and excess SiO₂. The liquid, which is silica-oversaturated, contains about 3.5% CO₂ and 10% H₂O. Back of that tetrahedron is the three-phase volume Fo+En+L; these vapor-absent liquids are also silica-oversaturated. Compositions on the plane En-H₂O-CO₂ or in the volume Fo-En-H₂O-CO₂ lying to the right of the stippled tetrahedron do not melt at 1375^oC.

At 1400° C, En and Fo are in equilibrium with a slightly more CO₂rich vapor and with a liquid that now lies exactly on the plane En-H₂O-CO₂ and is exactly silica-saturated. It is clear that the presence of CO₂, even in amounts sufficient to saturate the liquid and produce a free vapor phase, does not alter the silica-saturated or undersaturated nature of the melt. At 1450° C, En+Fo, both in equilibrium with a CO₂rich vapor and in the vapor-absent region, are in equilibrium with silica-undersaturated liquids. At 1325° C, all liquids, even those in equilibrium with Fo or Fo+V, are silica-oversaturated.

The assemblage Fo+En is a model peridotite. Figure 1 shows that the temperature at which that assemblage melts is a function of the $H_{2}O-CO_{2}$ content of the source region. The composition of the first-formed melt is a function of temperature (and indirectly, therefore, of $H_{2}O-CO_{2}$ ratio). Compositions containing more than 55 mole % CO_{2} will melt above 1400° C, and liquids are silica-undersaturated. These relationships hold whether or not the liquids are vapor-saturated.

The dominance of the melting behavior of enstatite in controlling composition of melts produced from peridotite is seen from experiments of Mysen (1973). He has found a change in composition of liquids produced by melting natural peridotite from silica-oversaturated to undersaturated at nearly the same CO_2 -H₂O ratio (0.55) as in the simple system discussed here.

Conclusions. Petrologic studies of kimberlite and related rocks have indicated that a CO2-H2O vapor phase was involved in their origin. This phase equilibria study indicates at least three reasons why such vapor associated with kimberlite and alkaline rocks should be CO2-rich: (1) The model system Fo-En-H20-CO2 indicates that in the presence of CO2 or CO2-H2O mixtures peridotite comprising the upper mantle melts at higher temperature. Although the presence of CO₂ has little effect on the silica-saturated or undersaturated nature of liquid produced by partial melting at any particular temperature, liquids produced at higher temperatures are less silica-saturated because of the "enstatite effect". Higher melting temperatures also produce a smaller degree of melting, at any point along the geotherm, yielding more alkalic melts. The effect of CO₂ is therefore the production of less silica-saturated, more alkalic, melts. There may be primordial areas of the mantle rich in CO2, or CO2-rich vapor may rise to certain localities. Mantle regions with vapor richer in H₂O would melt more completely to yield silicic magmas. (2) Silicate melts dissolve rather large amounts of H_2O at high pressures. The presence of a free vapor is unlikely. Melts dissolve much less CO2, so that the miscibility gap between melt and CO2-rich vapor is wider, and the presence of a free vapor is more likely.



(3) Vapor rich in CO_2 does not dissolve excess silica, unlike H_2O -rich vapor and may well dissolve excess alkalies. Such vapors would produce many alteration effects observed in kimberlite.



REFERENCES CITED

Boyd, F.R., and P.H. Nixon, 1973, <u>Carn. Inst. Wash. Y.B.</u>, in press. Dawson, J.B., 1971, <u>Earth-Science Reviews</u>, 7, 187-214.

Eggler, D.H., 1973, Carn. Inst. Wash. Y.B., in press.

Koster van Groos, A.F., and P.J. Wyllie, 1973, <u>Amer. Jour. Sci., 273</u>, 465-487.

Kushiro, I., H.S. Yoder, Jr., and M. Nishikawa, 1968, <u>Geol. Soc. Amer.</u> <u>Bull., 79</u>, 1685-1692.

Mysen, B.O., 1973, Carn. Inst. Wash. Y.B., in press.