THE PETROCHEMISTRY OF KIMBERLITE AUTOLITHS

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Diatreme facies of kimberlite frequently contain nucleated spheroidal bodies varying in size from a few millimetres to 70 millimetres in diameter. These autoliths consist of rock or mineral fragment nuclei which usually make up only a small volume percentage, encased in fine-grained kimberlitic material. It is thought that autoliths represent crystallisation of kimberlitic magma around a nucleus so that these bodies are more likely to reveal the matrix composition of kimberlites than the normally chaotic host kimberlite with its attendant xenolithic and xenocrystic material (Ferguson <u>et al</u>. in press). The nuclei of the autoliths comprise both mantle and crustal derived fragments. Olivine is the dominant mineral of the autoliths occurring as euhedral to subhedral phenocrysts 0.15 to 0.80 mm in diameter and having compositional range of Fo_{90-94} . Other possible primary phases include apatite, calcite and dolomite. Alteration reaction and minor accidental products include serpentine, ilmenite, phlogopite, spinels, perovskite and rutile.

Major element compositions were determined for a total of 26 autoliths from Wesselton Mine and various localities in Lesotho, and a cluster analysis program was used to compare the chemistry of these autoliths with 96 kimberlites and alkaline ultramafic rocks associated in space and time. It was found that the major element chemistry of the autoliths did not offer any unique composition but grouped with two of five varieties of kimberlite. Relative to the major kimberlite groups not clustering with autoliths the latter are enriched in TiO₂, P₂O₅, K₂O and MnO and impoverished in MgO. Further confirmation of the similarities between autoliths and kimberlites is given by their low K:Rb and Sr^{O7}:Sr^{O6} values. The Sr⁸⁷:Sr⁸⁶ ratio for Lesotho autoliths gives values between 0.7040 and 0.7045 (Hugh Allsopp, personal communication), which are comparable to the lowest values yet recorded for kimberlites (Berg and Allsopp, 1972).

On the CMAS tetrahedron (O'Hara, 1968) kimberlites lie on a welldefined olivine control line trending sub-parallel to the CAM plane consistent with melting of a four-phase garnet lherzolite mantle rock at depths equivalent to more than 40 Kb. Autoliths have compositions which plot near the extremity of the main kimberlite trend showing no major differences to most kimberlite other than having fractionated more olivine than the more primitive varieties. This supports the contention that the autoliths represent kimberlite magma which has only undergone olivine fractionation and has erupted from depth equivalents

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of greater than 40Kb. The alkaline ultramafic rocks that are associated in space and time with the kimberlites plot at the extremity of the kimberlitic trend with some degree of overlap indicating possible consanguinity (Ferguson <u>et al</u>, in press).

PICROILMENITES IN THE AUTOLITHS

Major element compositions have been determined in the electron microprobe for a total of 70 ilmenites included in the autoliths, and five autolith nuclei. The results of these analyses are summarized in Table 1, where they are compared with other relevant ilmenite analyses from the literature.

TABLE 1

COMPOSITIONS OF MAGNESIAN ILMENITES FROM AUTOLITHS, KIMBERLITE, AND ILMENITE-SILICATE NODULES (%)									
	1	2	3	4	5	6	7	8	
Ti0 ₂	52.24	53.98	50.98	47.27	52.63	56.40	56.80	53.01	
Cr ₂ 0 ₃	1.20	2.13	1.02	1.67	1.42	1.13	2.27	5.04	
*Fe203	9.20	6.81	10.26	14.41	0.33	1.91	0.57	5.20	
FeO	22.41	20.61	26.76	28.97	32.17	19.90	27.09	19.42	
MgO	13.64	15.10	10.71	7.60	12.02	20.23	13.19	15.74	
2. Aver 3. Aver 4. Aver 5. Prim 6. Secc 7. Secc 8. Ilme Isonvil	age Wes age Les age aut ary Liq ondary L ondary i enite in le, Ken	selton otho au olith n hobong iqhobon lmenite clusion tucky,	autolit tolith ucleus ilmenit g ilmen in gar in euh U.S.A.	h ilmen ilmenit e (Hagg ite (Ha net lhe edral o (Boyd a	ite e gerty, I ggerty, rzolite livine nd Nixo	n Press In Pre (Boyd crystal on, In P) and Nix from k ress).	on) imberlit	e,
The M greater th shows that ferric irc from the I ilmenites, MgO and Fe generally common. A ship betwe	IgO cont an ten these on than sotho and th 3 ⁺ . Th high, a lso of een Cr a	ents of weight ilmenit autolit ose fro e chrom nd valu interes nd Fe ³⁺	the au per cen es cont es occu hs appe m the W e conte es in e t is a	tolith t, and ain sig rring a ar to b esselto nts of xcess of fairly nd also	ilmenit inspect mifican s autho be trans on autol the aut of 3.0 p well de o noted	es are, ion of tly mor lith nu sitional iths, b colith i per cent eveloped by Boyd	with f the dat e magne clei. betwee oth wit lmenite Cr ₂ O ₃ invers and Ni	ew excep a in Tab sia and The ilme n the nu h respec s are are not e relati xon (In	tions le l less nites cleus t to un- on-

Of particular importance, however, is the marked similarity in composition between the autolith ilmenites and secondary ilmenites from the Lighobong kimberlite (Haggerty, In Press) and from a garnet

Press) for ilmenites intergrown with various silicate minerals.

lherzolite from Monastery Mine (Boyd and Nixon, In Press) (Table 1).

Haggerty (op. cit.) has described mantled sequences of ilmenite and spinel on picroilmenite in the Lighobong kimberlite, and he has shown that the zoning with respect to Mg/Mg + Fe in these sequences is inverted, and that secondary ilmenites thus produced are conspicuously more magnesian than the primary ilmenites from which they were derived. Like the autolith ilmenites, these secondary ilmenites appear to be enriched in Cr, and depleted in Fe³⁺ relative to ilmenites intergrown with various silicate minerals (Boyd and Nixon, op. cit.).

These observations support Boyd and Nixon's suggestion that these magnesian ilmenites are of a later generation than the more Fe-rich variety, of which the autolith nuclei are excellent examples. From the work of Boyd and Nixon (op. cit.) and Mitchell (1973) it is evident that much of the ilmenite now found in kimberlite was formed at great depth, possibly in the low-velocity zone (Boyd and Nixon, op. cit.), and possibly also before the generation of diamond (Mitchell, op cit.).

Clearly, therefore, kimberlitic ilmenites have formed over wide ranges of pressure, temperature and oxygen fugacity. It is also evident that although the autoliths contain occasional low-magnesian ilmenites, the majority are the more magnesian variety, and the abundance of these ilmenites, together with secondary spinels in the autoliths provide strong evidence for the existence of highly reactive kimberlite liquids prior to and during kimberlite eruption.

It is concluded that autoliths represent kimberlite matrix having precipitated around solid nuclei during pipe development of kimberlite intrusions. They have only been subjected to olivine fractionation and have erupted from depths in excess of 100 km.

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