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# HIGH-Mg CARBONATITIC HDFs, KIMBERLITES AND THE SCLM

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# INTRODUCTION

The story of diamonds is interwoven with that of kimberlitic volcanism (Gurney et al., 2010), but the genetic relationship between them has been debated ever since kimberlite was first identified. At first, diamonds were believed to be phenocrysts from early crystallization phases of the kimberlitic magma (Dawson, 1971; Harte, 1980). However, evidence for deep lithospheric pressures for the entrapment of inclusions (see reviews by Meyer, 1987; Stachel and Harris, 2008), the occurrence of diamonds in lamproites and mantle xenoliths, and the recognition that most diamonds are much older than their host kimberlite (Richardson et al., 1984; Gurney et al., 2010) has led to a consensus that most diamonds are xenocrysts in kimberlites. The case for a kimberlitic origin was still raised for fibrous diamonds (Javoy et al., 1984; Boyd et al. 1987). However, the compositions of the high-density fluids (HDFs) they have trapped indicate that they are xenocrysts as well. Navon et al. (1988) pointed out a broad similarity in majorelement composition between the fluids trapped in fibrous diamonds and the bulk compositions of Group II kimberlites, and suggested a genetic link at depth. Α general similarity between the trace-element abundance patterns of HDFs and kimberlites was noted first by Akagi and Masuda (1988) and later by Schrauder et al. (1996) Zedgenizov et al. (2007), Tomlinson et al. (2009) and Rege et al. (2010), but no attempt has been made to test this idea quantitatively.

In the present work we compare the trace-element data of high-Mg carbonatitic HDFs from Kankan, Guinea, and Udachnaya, Sibiria, with that of kimberlites and consider models for a possible genetic links between them.

#### **HIGH-Mg CARBONATITIC HDFs**

The high-Mg carbonatitic HDFs are characterized by very high MgO and CaO content. The  $SiO_2$  and  $Al_2O_3$  content are low and the  $K_2O$  content is variable. It is very high in Siberian HDFs and lower in Guinean ones (Klein-BenDavid et al., 2009a; Weiss et al., 2009). All are highly enriched in the most incompatible trace-elements, with levels of up to a

few thousand times the primitive-mantle values (Figure 1). The Kankan high-Mg carbonatitic HDFs show depletion of K, Rb, Cs, Nb and Ta and enrichment in Ba, Th, U and LREE relative to the Siberian HDFs. These differences in the highly incompatible elements (Cs–Sr) define two characteristic patterns. Ratios such as La/Nb, Th/Nb, Nb/Rb and Sr/Rb highlight these differences and can be used to distinguish between the two suites. Elements that are more compatible than Sr show similar characteristics and define overlapping patterns (Figure 1).



Figure 1: Trace-element patterns of high-Mg carbonatitic HDFs. The different symbols represent the trapped HDFs in six Udachnaya diamonds while the shaded area represents the range for the HDFs in three Kankan diamonds. The uncertainties on the values of Yb and Lu are large and they can be regarded as qualitative only. Primitive-mantle values are from McDonough and Sun (1995).

#### THE KIMBERLITE CONNECTION

The major-element compositions of the high-Mg HDFs (Klein-BenDavid et al., 2009a; Weiss et al., 2009; Zedgenizov et al., 2007, 2009) resemble experimentally-produced near-solidus melts of carbonate-bearing peridotite (Brey et al., 2008; Dasgupta and Hirschmann, 2007). The high-pressure experimental results of Dalton and Presnall (1998a,b) in the system CaO-MgO-Al<sub>2</sub>O<sub>3</sub>-SiO<sub>2</sub>-CO<sub>2</sub> connect the high-Mg carbonatitic HDFs and the near-solidus melts of carbonated peridotite with both Group I

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and Group II kimberlites (Figure 2). This suggests that with continuous heating, an initial high-Mg carbonatitic fluid becomes poorer in CaO and richer in SiO<sub>2</sub>,  $Al_2O_3$  and MgO and evolves toward a kimberlitic melt.

The trace-element patterns of Udachnaya high-Mg HDFs exhibit close similarity to those of Group I kimberlites from Africa, Canada, Siberia, Greenland, Finland, India and Brazil (Figure 3). Compared to Group II kimberlites the Udachnaya high-Mg HDFs reveal higher Nb and Ta relative to Ba, Th, U, La and Ce. This can be seen best using La/Nb *vs* Th/Nb diagram. The Udachnaya HDFs fall in the field of Group I kimberlites, whereas the Kankan high-Mg HDFs, with their lower Nb and Ta and higher correlative ratios, fall close to Group II kimberlites (Figure 3).

Overall, the major- and trace-element similarities suggest a genetic link between the high-Mg HDFs from Udachnaya and Group I kimberlites and between Kankan HDFs and Group II kimberlites.



**Figure 2:** Major-element plots comparing HDFs and kimberlites with near-solidus melts of carbonated peridotite. The trend of kimberlites to much higher MgO/CaO reflects assimilation of olivine.

## **CRYSTALLIZING THE ASCENDING KIMBERLITE**

Cooling of kimberlitic melt and crystallization of olivine, garnet, pyroxene, carbonate and ilmenite may lead to the formation of high-Mg carbonatitic residual fluids. Figure 4 shows the trace-element pattern of the residual fluid after 80% crystallization. For most elements, the overall pattern of the resulting fluid is very similar to that of the measured high-Mg HDF. It is harder to reconcile the low initial K<sub>2</sub>O content of the kimberlite with the high K<sub>2</sub>O of the Siberian HDFs, or their similar water content, with the predictions of



**Figure 3:** Trace-element characteristics of high-Mg carbonatitic HDFs and kimberlites. (a) The average trace-element pattern of high-Mg carbonatitic HDFs in diamonds from Udachnaya and Kankan, worldwide average Group I kimberlite pattern (green squares) as determined by averaging the compositions of Group I kimberlites from Africa, Canada, Siberia, Greenland, Finland, India and Brazil (thin black lines within the light green area) and the average pattern of South African Group II kimberlites. (b) La/Nb vs Th/Nb values in Group I and Group II kimberlite compared with high-Mg carbonatitic HDFs in Kankan and Udachnaya diamonds. See Weiss et al. (2011) for data sources.



**Figure 4:** Crystallization of kimberlite to form high-Mg HDF. Ti concentration is based on the mass balance calculation of the cumulate. The values for the average Group I kimberlite are as in Figure 3.

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a fractional crystallization model. However, these discrepancies cannot fully dismiss such a scenario, as the  $K_2O$  content of the kimberlite may have been higher at depth (Kamenetsky et al., 2004) and the water discrepancy may have its origin in the near-surface modification of the kimberlite (Sheppard and Dawson, 1975; Bell et al., 2004).

# PARTIAL MELTING OF SIMILAR SOURCES

Another way to test the linkage between kimberlites and high-Mg HDFs is by modeling them as product of melting of a common source. We calculated the trace element peridotite (66% concentrations of olivine, 26% orthopyroxene, 3% clinopyroxene and 5% garnet) in equilibrium with the Udachnaya high-Mg HDFs, average Udachnaya kimberlite, an average South African Group I kimberlite and a worldwide average of Group I kimberlites (Figure 5a). All the calculated sources have similar patterns. The closest match of these calculated sources is obtained using partial melting of F=3% for the Udachnaya kimberlite, 2% for the South African and worldwide average Group I kimberlite and 0.2% for the high-Mg HDFs. The similarity in trace-element patterns between the calculated sources of Group I kimberlites and the Udachnaya high-Mg HDFs, combined with experimental studies that fit their major-element compositions (Dalton and Presnall, 1998a,b; Brey et al., 2008), strongly suggest that the two enriched melts can be formed by different degrees of melting of the same (or very similar) mantle source. The <sup>87</sup>Sr/<sup>86</sup>Sr ratios of the trapped high-Mg HDFs in four of the Udachnaya diamonds (Klein-BenDavid et al., 2009b) vary between 0.7043 and 0.7053, well within the range of Group I kimberlites (0.7030-0.7055; Smith, 1983; Becker and le Roex, 2006 and references therein). This correspondence in <sup>87</sup>Sr/<sup>86</sup>Sr ratios provides further support for the suggested connection between Udachnaya high-Mg HDFs and Group I kimberlites.

High-Mg HDFs trapped in diamonds from Kankan, Guinea have lower levels of alkalis and higher contents of LREE, Ba, Th and U than the Udachnaya high-Mg HDFs (Figure 1 and 3). This may reflect smaller degrees of melting compared to the Siberian HDFs and retention of phlogopite in their source rocks. Similarly, the composition of Group II kimberlites argues for phlogopite in their source (Mitchell, 1995), and they represent lower degrees of melting than Group I kimberlites (Becker and le Roex, 2006). When calculating the source compositions, a good fit is obtained if the Kankan HDFs represent F=0.1% and are in equilibrium with a source that carries 0.3% phlogopite, while Group II kimberlites are produced by 1% melting in equilibrium with a source with 0.1% phlogopite (Figure 5b). The lower amount of phlogopite in equilibrium with Group II kimberlite is consistent with their relatively higher degree of melting, their higher Al<sub>2</sub>O<sub>3</sub> and SiO<sub>2</sub> compared to the Kankan high-Mg HDFs (Figure 2), and their higher K<sub>2</sub>O contents compared to Group I kimberlites

(Mitchell, 1995; Becker and le Roex, 2006). Here as well, farther support is provided by the high <sup>87</sup>Sr/<sup>86</sup>Sr ratio (0.7079) in a single Kankan diamond, ON-KAN-386 (Klein-BenDavid unpublished data), well outside the range of Group I kimberlites but within the range of Group II kimberlites (<sup>87</sup>Sr/<sup>86</sup>Sr=0.7072–0.7083; Smith et al., 1983; Becker and le Roex, 2006 and references therein).



**Figure 5:** Calculated source trace-element composition for high-Mg HDFs and kimberlites. (a) Udachnaya high-Mg HDFs and Group I kimberlites and their calculated source patterns. (b) Kankan high-Mg HDFs and average South African Group II kimberlites and their calculated sources. Mineral/carbonatitic-melt partition coefficients for olivine, orthopyroxene, clinopyroxene and garnet are from Dasgupta et al. (2009) and for phlogopite from Sweeney et al. (1995) and Schmidt et al. (1999).

## DIAMOND-FORMING HIGH-Mg HDFs AND KIMBERLITES AS METASOMATIC AGENTS

It is now widely accepted that the sub-continental lithospheric mantle (SCLM) was strongly depleted through extensive extraction of mafic melts and that it was later refertilised by interaction with trace element-rich mantlederived melts and fluids (Erlank et al., 1987). In spite of the diverse patterns of individual xenoliths, the average trace-element compositions of Archaean xenoliths from the Siberian, Canadian, North Atlantic and the Kaapvaal cratons are highly comparable (Schmidberger and Francis, 2001; Gre'goire et al., 2003; Ionov et al., 2010; Simon et al., 2007; Wittig et al., 2008). The similarity is surprising, as the lithosphere in the various localities are separated both



physically and by the age of formation. All these averages are different from the calculated average of basalt-hosted peridotite xenoliths from the post-Archaean SCLM (McDonough, 1990).

The trace-element patterns of Udachnaya high-Mg carbonatitic HDFs and Group I kimberlites are similar in shape to the pattern of the average post-Archaean xenolith rather than to the average of the Archaean ones (Figure 6a). This is remarkable since the HDF-bearing diamonds and many kimberlites are found on Archaean cratons. Figure 6b shows that the similarity in shape can be explained by a simple mixing model with addition of 2.5% of average Group I kimberlitic magma or 0.5% of the Udachnaya high-Mg carbonatitic HDF to a depleted SCLM (modeled as MORB source or a more depleted one). The question remains as to why the average post-Archaean, rather than the Archaean, xenoliths have trace element patterns that are similar to those of kimberlitic magmas or high-Mg carbonatitic HDFs.

The average compositions of different regions of the SCLM may actually be referred to as the average enrichment, due to different metasomatic events that the SCLM experienced between the time of its depletion and its sampling as xenoliths. Thus, the trace-element similarity of the average post-Archaean xenoliths may be the result of metasomatism mostly by kimberlites, high-Mg HDFs and other melts with similar patterns. Some of the Archaean xenoliths have enriched trace-element patterns that are comparable to the average of the post-Archaean ones, suggesting that they as well were metasomatically influenced by such melts. On the other hand, many Archaean xenoliths and their cratonicaverage compositions are characterized by different patterns, suggesting that fluids other than kimberlites or high-Mg HDFs were metasomaticaly enriching the Archaean SCLM. Indeed, other diamond-forming HDFs with different compositions and patterns have been found (10IKC abstract No. 174). Mixing of these HDFs with a depleted SCLM source explains some of the different traceelement characteristics, as sampled by Archaean metasomatized xenoliths. The similarity of the average pattern of xenoliths from various Archaean cratons suggests that long integration of interactions with different metasomatic agents leave broadly similar, large-scale, enriched litospheric reservoirs.

# CONCLUSIONS

The similar trace-element patterns of kimberlites and HDFs strongly support a common source for both. The patterns and element ratios also suggest that the Udachnaya HDFs are closer to Group I kimberlite, while the Kankan HDFs are associated with Group II. These associations are reflected in their respective Sr-isotope compositions. The close relation between HDFs that were trapped at depth and kimberlites sampled at the surface, places constraints on the chemical composition of kimberlitic magmas at depth and

indicates that their incompatible-element compositions did not change much *en route*.

The similarity of the trace-element patterns of both HDFs and kimberlites to that of the post-Archaean SCLM may reflect the interaction of such melts with the lithospheric keel, its melting to produce HDF and/or kimberlites or melting of deeper sources that led to formation of HDFs and kimberlite and to widespread metasomatism of the SCLM.



Figure 6: (a) Trace-element patterns of the Udachnaya high-Mg HDFs, worldwide average Group I kimberlites (from Figure 3) and the average composition of the Archaean and post Archaean SCLM. (b) The enriched post-Archaean SCLM is closely fitted by mixing either kimberlitic melt or high-Mg HDF with depleted mantle source, represented by the average depleted MORB mantle (DMM) composition (other, more depleted composition would yield similar results).

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