

## Formation and temporal evolution of the Venetian Kaapvaal lithosphere

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The South African Venetia kimberlite pipes are situated in the metamorphic crust of the Limpopo Mobile Belt. Seismic tomography (James et al., 2001) suggests that the metamorphic crust is underlain by mantle lithosphere similar to the Zimbabwe and Kaapvaal Cratons. Furthermore, U/Pb studies on zircons from the Venetia region suggest crustal protolith ages as old as 3.65 Ga (e.g., Zeh et al., 2007), in close similarity to Archaean Cratons.

Xenoliths from the Venetia kimberlite pipes are investigated in order to establish if silica enrichment of the sub-continental lithospheric mantle (SCLM) beneath the Archaean Kaapvaal craton originated prior to kimberlite emplacement at ~533 Ma. Notably, Venetia xenoliths were unaffected by the Karoo flood basalt event and Group II type kimberlite magmatism and diamond inclusion data argue for extreme major element depletion in the SCLM beneath Venetia (Viljoen et al., 1999). Venetian diamond inclusion trace element data (Stachel et al., 2004) and previous major and trace element studies on Venetian mantle xenoliths (Stiefenhofer et al., 1999; Barton and Gerya, 2003) suggest that cryptic (silicate melt dominated) metasomatism occurred. Rhenium depletion ages for 7 mantle xenoliths yield two age populations of ~1.4 Ga and >2.5 Ga (Carlson et al., 1999), apparently unrelated to the regional Limpopo tectono-magmatic event. Estimates of the time of diamond formation range from less than 50 Myr prior to eruption to an older generation that formed at ~2.5 Ga, with eclogitic inclusions formed at 2.05 Ga (Burgess et al., 2004; Richardson et al., 2006; Richardson and Shirey, 2008).

550 peridotite xenoliths were studied in the field; 55% harzburgites, 37% lherzolites and 8% dunites/pyroxenites, all were heavily altered. To avoid the effects of metasomatic events and severe alteration, 35 relatively fresh harzburgites were selected for a detailed petrological and geochemical study. Petrographically, three groups of xenoliths were recognised. Group I consists of 6 cpx-free garnet harzburgites. One has an equant coarse texture, the other five are porphyroblastic. Garnet compositions are sub-calcic and olivine has Mg# of 91.7-93.4. Mineral major element data suggest that garnet, olivine and opx are in major element equilibrium, except for one sample that may have suffered from slight metasomatism by an asthenosphere-derived melt as suggested by variably TiO<sub>2</sub> enriched garnet and opx

Mg# from 90.8 to 93.4. Group II includes 14 spinel lherzolites of which 8 have coexisting garnet. They have equant coarse textures, although opx is sometimes lobate and garnets are often interlobate to vermicular forming (linear) clusters. Cpx (<2%) often occurs as texturally equilibrated grains. Mineral chemistry is very homogeneous with olivine Mg# of 92.2-92.7 and all minerals appear to be in major element equilibrium (Fig. 1). Group III includes 15 garnet lherzolites with porphyroblastic textures covering a range of compositions, reflected by olivine Mg# of 88.7-92.3. About half of the samples contain ~2% poikiloblastic cpx. Mineral major element data suggest that minerals in all but two samples are in major element equilibrium. Opx abundances among the three groups vary from <10% in some group I samples to about 25%, with most samples having ~20% opx. This suggests that SiO<sub>2</sub> enrichment of the SCLM had occurred prior to entrainment of Venetian mantle xenoliths at ~533 Ma, although the enrichment seems less extreme than beneath the Kaapvaal Craton (e.g., Pearson and Wittig, in press).

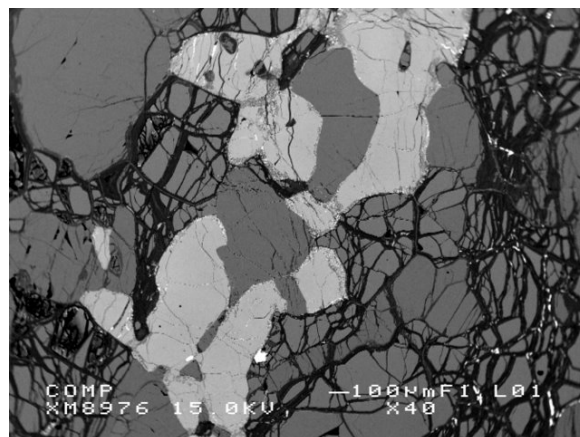


Fig. 1. Back scatter electron image of a garnet-bearing group II sample. Garnet (light grey reflection) is not fully texturally equilibrated and often occurs clustered having interlobate to vermicular textures.

Pressure and temperature calculations (mostly from groups I and III) range from 5.5 GPa at 1270°C to 6.4 GPa at 1451°C. The exceptions are an equant coarse garnet harzburgite (5.2 GPa at 1126°C) and two group II samples (lowest PT of 2.7 GPa at 826°C). All samples plot close to a 43 mW m<sup>-2</sup> geotherm.

Following the work of Bernstein et al. (2007), olivine

Mg# suggest that melt percentages for group I samples vary from ~28% to ~53%, whereas estimates for group II samples display a smaller range from 34% to 42% (Fig. 2). Group III samples have lower melt percentages between 21 and 38% excluding the two samples with the lowest Mg# (90.0 and 88.7) close to or lower than assumed fertile mantle, suggesting modification by metasomatism.

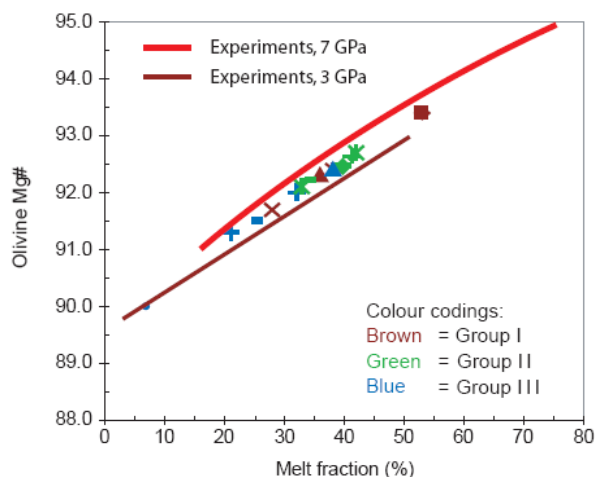


Fig. 2. Olivine Mg# versus melt fraction. Venetia melt fractions vary from ~21 to ~53%, with the highest estimates for group I samples (brown symbols) and group III samples (blue symbols) generally on the lower side. Different symbol shapes refer to different samples (same in each figure). Distinction between high (7 GPa) and low (3 GPa) pressure melting is not possible. (After Pearson and Wittig, in press.)

In situ LA-ICPMS on 8 samples reveal large heterogeneity, e.g., garnet  $[La]_N$  varies from 0.05 to 1.9 and  $[Lu]_N$  from 0.97 to 11.8. LREE in opx are also highly variable with  $[La]_N$  between 0.03 and 2.7. Distinction between the three groups for garnet trace elements is not as clear as for major element chemistry, while opx trace element compositions show no systematic variation. Garnet REE compositions vary from LREE depletion ( $[Ce/Yb]_N = 0.19-0.50$ ) in group III to strongly sinusoidal patterns in group I, while group II samples are intermediate. Sinusoidal patterns can occur with tight peaks at  $[Sm]_N$  or  $[Eu]_N \sim 4.5$  or with wider peaks shifted to lighter REE compositions ( $[Ce]_N$  or  $[Pr]_N$  about 6.5 to 9.0). The latter resemble fluid-dominated metasomatism (Stachel et al., 2004) in contrast to generally melt metasomatised Venetian mantle xenoliths (Stiefenhofer et al., 1999). Garnet diamond inclusions have dominantly sinusoidal  $[REE]_N$  patterns (Stachel et al., 2004), but also vary considerably in composition. Whole rock HREE contents were reconstructed from garnet, opx and (if present) cpx trace element data and modal abundances and compared to simple melt depletion modelling following Hellebrand et al. (2002) (Fig. 3). Group I and II samples can be modelled by either ~15% melting in the spinel field or combined 20% garnet field melting followed by ~15% spinel field melting. The reconstructed whole rock REE patterns do not distinguish between ~15% spinel field melting alone

and combined garnet+spinel field melting, but the total of ~35% for the latter scenario closely resembles melt extraction estimates derived from olivine compositions. Group III samples require less melting (as low as ~10%), but again dominantly in the absence of garnet as suggested by their low HREE contents ( $[Lu]_N$  in all samples varies from 0.8 to 0.05). The conclusion that ~15% spinel field melting is required for all samples implies later subduction to transport samples back to garnet stability field depths.

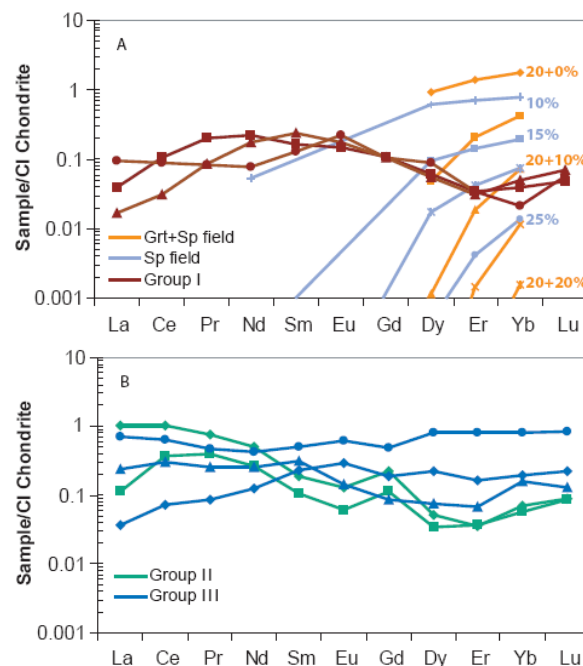


Fig. 3. Reconstructed whole rock  $[REE]_N$  patterns. Melting models for HREE concentrations are shown in (a) together with patterns for samples from group I. Patterns for samples from groups II and III are shown in (b). The low reconstructed whole rock HREE<sub>N</sub> concentrations for all samples indicate melting in the spinel field (minimum ~15% for all but 1 sample).

REE partition coefficients ( $K_D$ ) calculated between garnet and opx suggest that only one garnet lherzolite (B00-45 from group III) of the 8 samples is within error of equilibrium  $K_D$  from the literature. B00-45 was also the only sample with equilibrium gt/cpx  $K_D$ . Despite the inter-mineral trace element disequilibrium, hypothetical melts in equilibrium with garnet and cpx have been calculated from mineral trace element data and literature mineral/melt  $K_D$ . Constraints can be placed on the source of metasomatic agents by combining LILE, HFSE, LREE and HREE compositions of hypothetical melts and comparing these to true (natural) melt compositions. Calculated  $[Nb/Sr]_{PM}$  and  $[Ce/Yb]_{PM}$  in the hypothetical melts suggest that garnet formation is related to a mixture of metasomatism by hydrous fluids with elevated LREE/HREE (stage 1) followed by kimberlite-like metasomatism (stage 2; see Fig. 4). Calculated melts in equilibrium with cpx from (gt/cpx equilibrated) sample B00-45 plot within error of kimberlite and garnet-derived melts. A second melt calculated in equilibrium

with cpx from a group II sample plots close to the hydrous fluid field. This suggests cpx may have crystallised from hydrous fluids during stage 1 metasomatism. Alternatively, it is speculated that diffusive fractionation may have resulted in lower Nb/Sr and higher Ce/Yb than expected from partitioning behaviour.

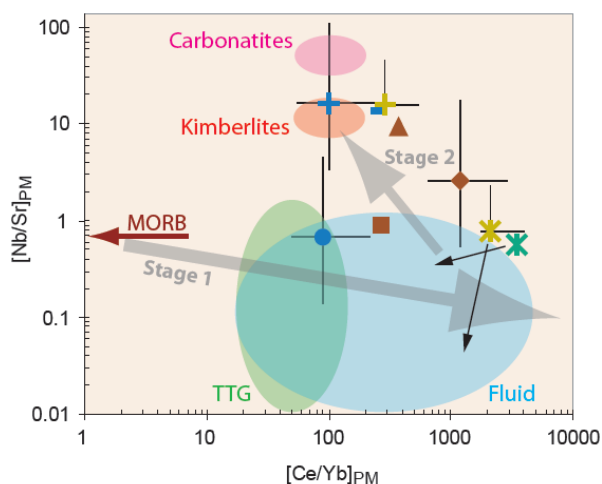


Fig. 4. Primitive mantle normalised [Nb/Sr] versus [Ce/Yb] in melts. Nb (HFSE) and Sr (LILE) were chosen because of relatively large concentration differences in the melts from different tectonic environments. Shaded regions represent natural compositions for carbonatitic (pink), kimberlitic (red), subduction zone siliceous melts (e.g., TTG, melts in equilibrium with eclogite; green) and hydrous fluids (blue). Yellow symbols are calculated melts in equilibrium with cpx (plus symbol from a group III, asterisk from a group II sample), while green (group I), brown (group II) and blue (group III) symbols are hypothetical melts in equilibrium with garnet. Error bars on calculated melt compositions (not on all data points for clarity) relate to literature variation in mineral/melt partition coefficients, while the small black arrows show the magnitude and direction of the change in melt compositions calculated with mineral/fluid instead of mineral/melt  $K_D$ .

In conclusion,  $\text{SiO}_2$  enrichment of the Kaapvaal Craton occurred prior to ~533 Ma. Venetian mantle xenoliths experienced on average ~15% melting in the spinel stability field, implying later tectonic subduction to transport them to garnet stability fields. Metasomatism occurred in at least two stages, first by hydrous fluids and subsequently by kimberlite-like melts. From diamond inclusions it can be inferred that the first stage took place after subduction had transported the samples to garnet stability fields and that the hydrous fluid stage of metasomatism is likely to be related to diamond formation. Ongoing Nd-Hf mineral isotope data will be used to deduce the timing of the above events.

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